

Modeling “Modern” Mantle Composition

- Models for the BSE (bulk silicate Earth) and the “upper” & “lower” mantle
- Uncertainty in models
- model of “lower” mantle composition
- Mass balance and implications for the composition of the BSE

Rick Arevalo (NASA/Goddard)

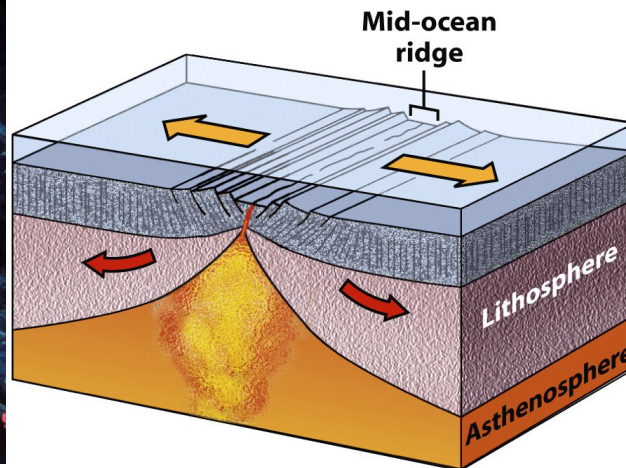
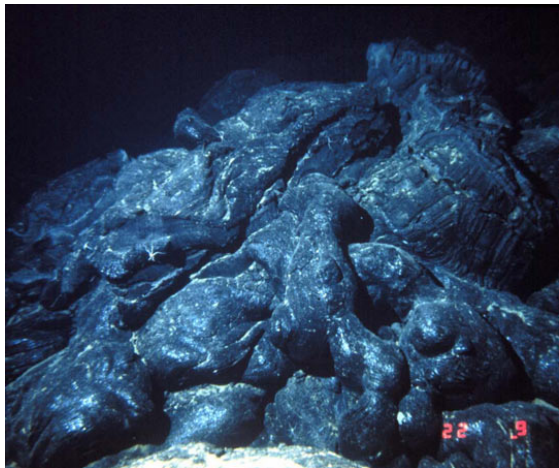
Bill McDonough (Univ of Maryland)



Why study the composition of MORB

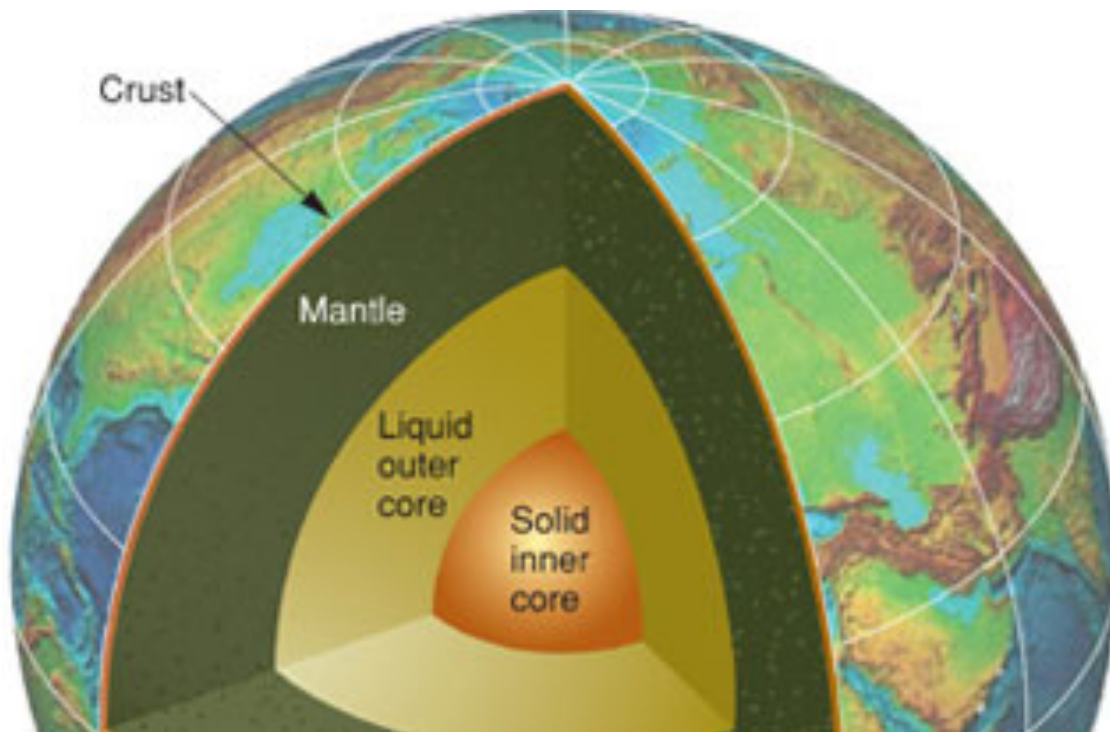
- Insights into the composition of the bulk Earth
- Understand differentiation of the BSE
(BSE : Bulk Silicate Earth = crust + mantle)
- Identify the distribution of elements in the Earth

Divergent Plate Boundary



The Mantle: it is the source of basalts

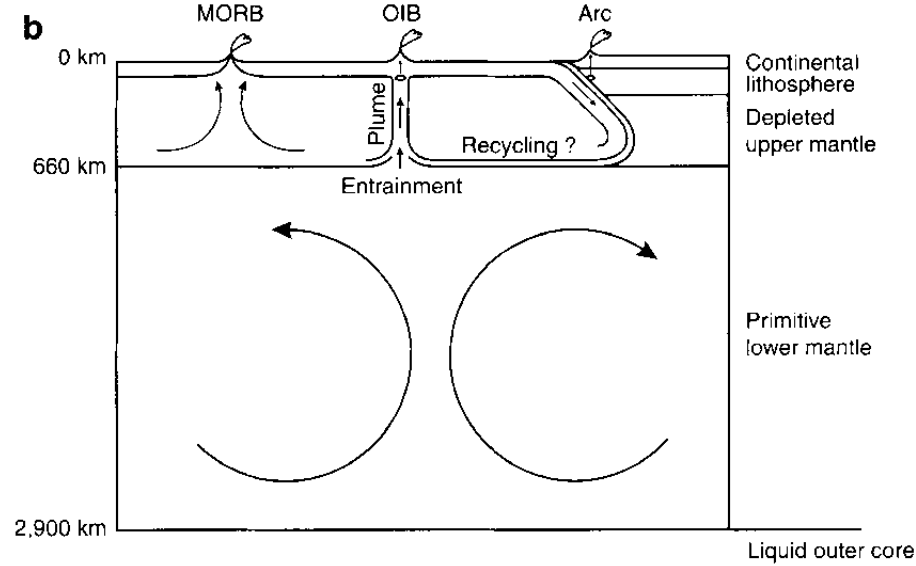
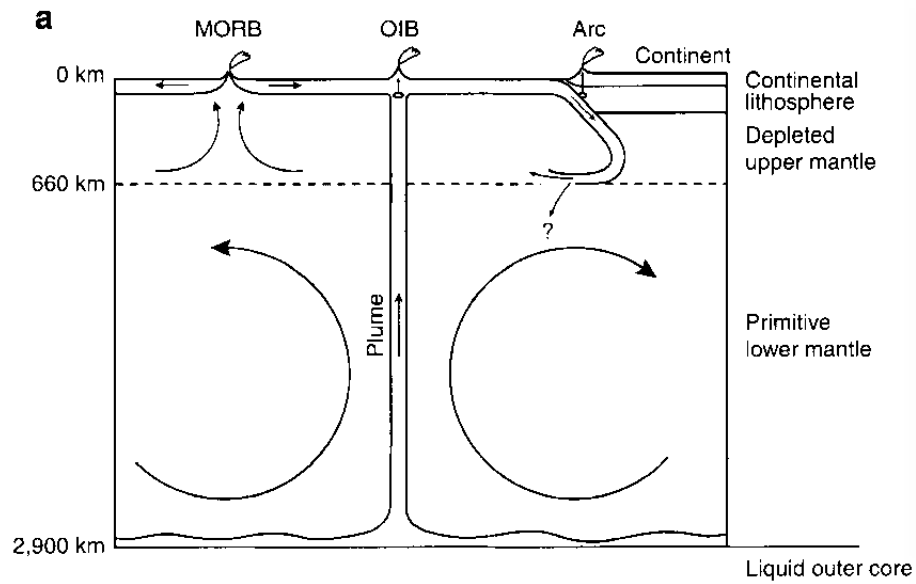
- 2900 km deep
- Surfaced by ~40km Continental or ~7km Oceanic Crust
- Surface potential temperature ~1550 K
- Core-Mantle Boundary (CMB) temperature 4000-5000 K



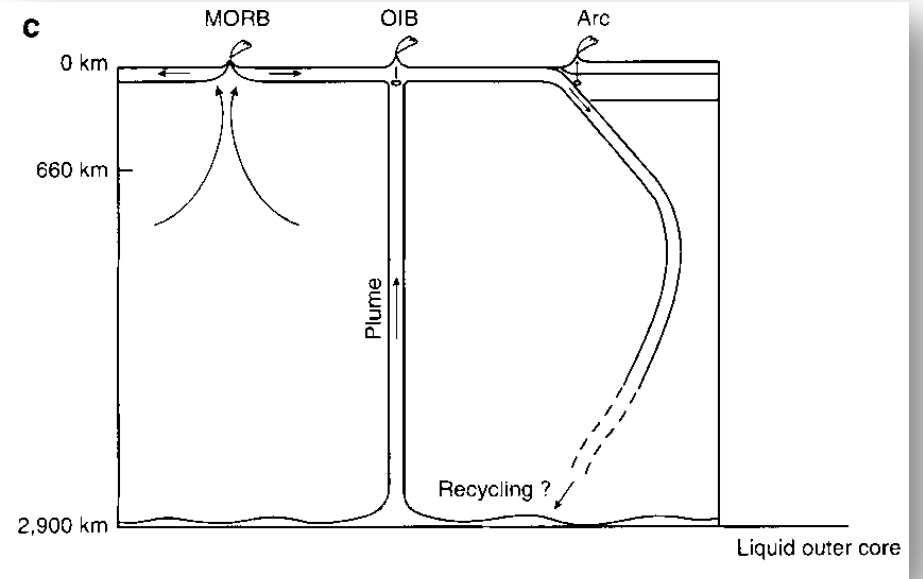
MORB Mantle

- Depth/Volume ?
- Top of mantle
- Residua from production of Continental Crust
- Recorder of convection efficiency

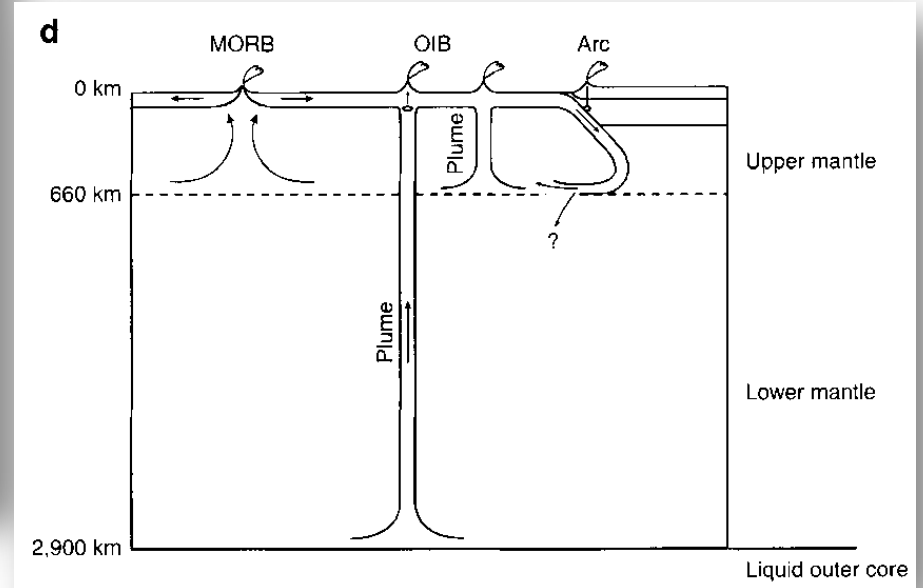
Convictional perspective



Non-layered model

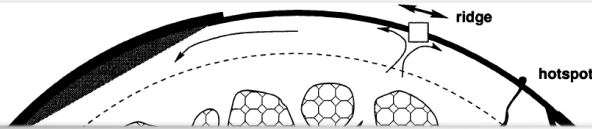


Non-layered model

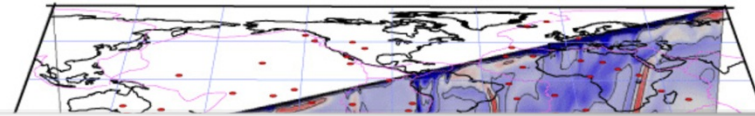


The simplest mantle models assume two distinct reservoirs: “upper” and “lower” mantle source regions.

BECKER et al. (1999) EPSL



McNamara and Zhong (2005) Nat

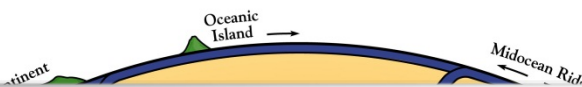


Assuming a two-component mantle:

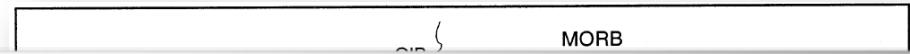
$$M^{BSE} \cdot X_i^{BSE} = M^{CC} \cdot X_i^{CC} + M^{UM} \cdot X_i^{UM} + M^{LM} \cdot X_i^{LM} \text{ or,}$$

Atoms of element i in BSE = Atoms i in CC + UM + LM

Kellogg et al. (1999) Sci



Morgan and Morgan (1999) EPSL



- ✓ **Masses of the BSE and CC are known**
- ✓ **Composition of the CC is well constrained**
- ✓ **Various models of BSE composition (discussed next)**
- ✓ **HERE, we focus on X_i^{LM} , and by extension M^{UM} and M^{LM}**

U content of BSE models

- Nucleosynthesis: U/Si and Th/Si production probability
- Solar photosphere: matches C1 carbonaceous chondrites
- Estimate from Chondrites: ~11ppb planet (16 ppb in BSE)
- Heat flow: secular cooling vs radiogenic contribution... ?
- Modeling composition: which chondrite should we use?

A brief (albeit biased) history of U estimates in BSE:

- | | |
|--------------------------------------|--|
| •Urey (56) 16 ppb | Turcotte & Schubert (82; 03) 31 ppb |
| •Wasserburg et al (63) 33 ppb | Hart & Zindler (86) 20.8 ppb |
| •Ganapathy & Anders (74) 18 ppb | McDonough & Sun (95) 20 ppb ± 20% |
| •Ringwood (75) 20 ppb | Allegre et al (95) 21 ppb |
| •Jagoutz et al (79) 26 ppb | Palme & O'Neill (03) 22 ppb ± 15% |
| •Schubert et al (80) 31 ppb | Lyubetskaya & Korenaga (05) 17 ppb ± 17% |
| •Davies (80) 12-23 ppb | O'Neill & Palme (08) 10 ppb |
| •Wanke (81) 21 ppb | Javoy et al (10) 12 ppb |

Disagreement with “chondritic” Earth Models

Murakami et al (May - 2012, *Nature*): “...the lower mantle is enriched in silicon ... consistent with the [CI] **chondritic Earth model**.”

Campbell and O’Neill (March - 2012, *Nature*): “Evidence **against a chondritic Earth**”

Zhang et al (March - 2012, *Nature Geoscience*): The Ti isotopic composition of the **Earth and Moon overlaps that of enstatite chondrites**.

Fitoussi and Bourdon (March - 2012, *Science*): “Si isotopes support the conclusion that **Earth was not built solely from enstatite chondrites**.”

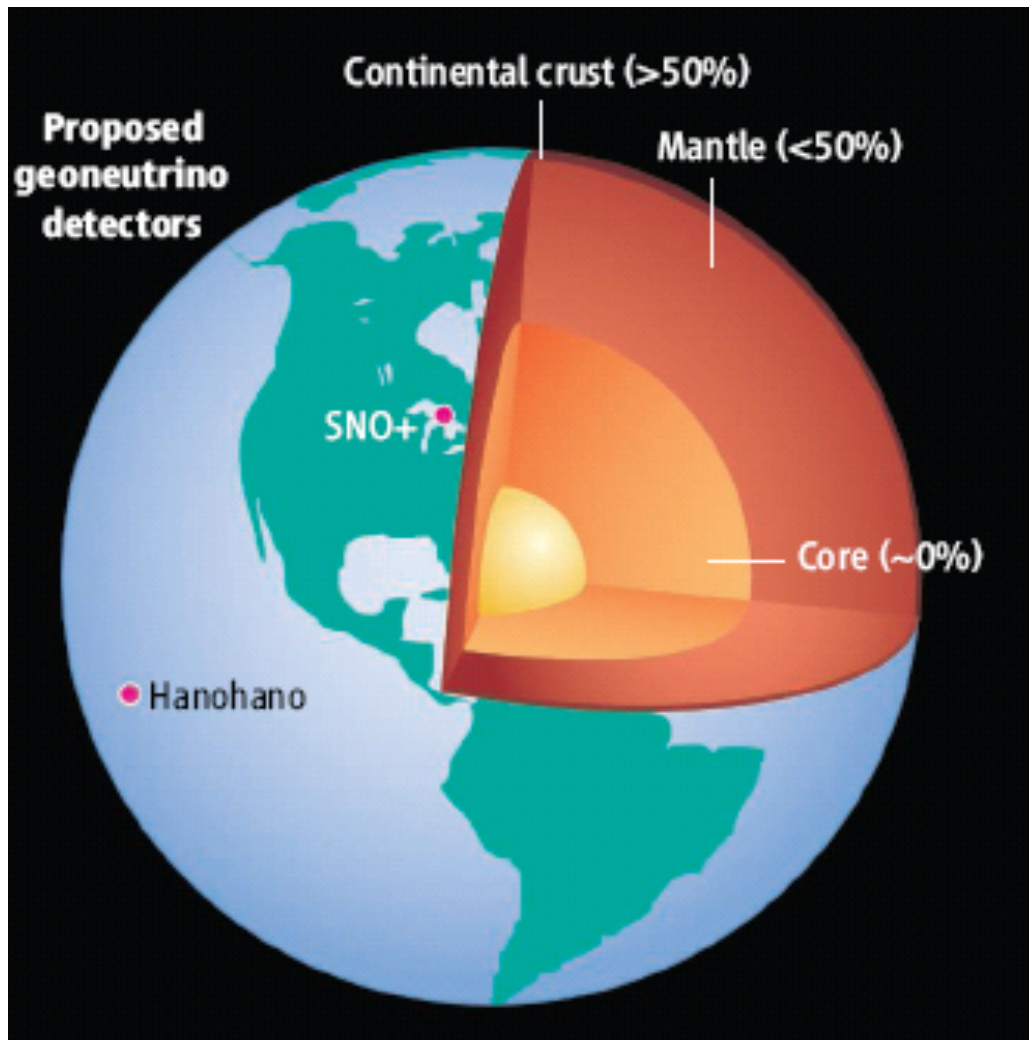
Warren (Nov - 2011, *EPSL*): “Among known chondrite groups, **EH yields a relatively close fit to the stable-isotopic composition of Earth**.”

- Compositional models differ widely, implying a **factor of three difference** in the U & Th abundances of the Earth



U in the Earth:

“Differentiation”



~13 ng/g U in the Earth

Metallic sphere (core)
<<<1 ng/g U

Silicate sphere
20* ng/g U

*O'Neill & Palme (2008) 10 ng/g

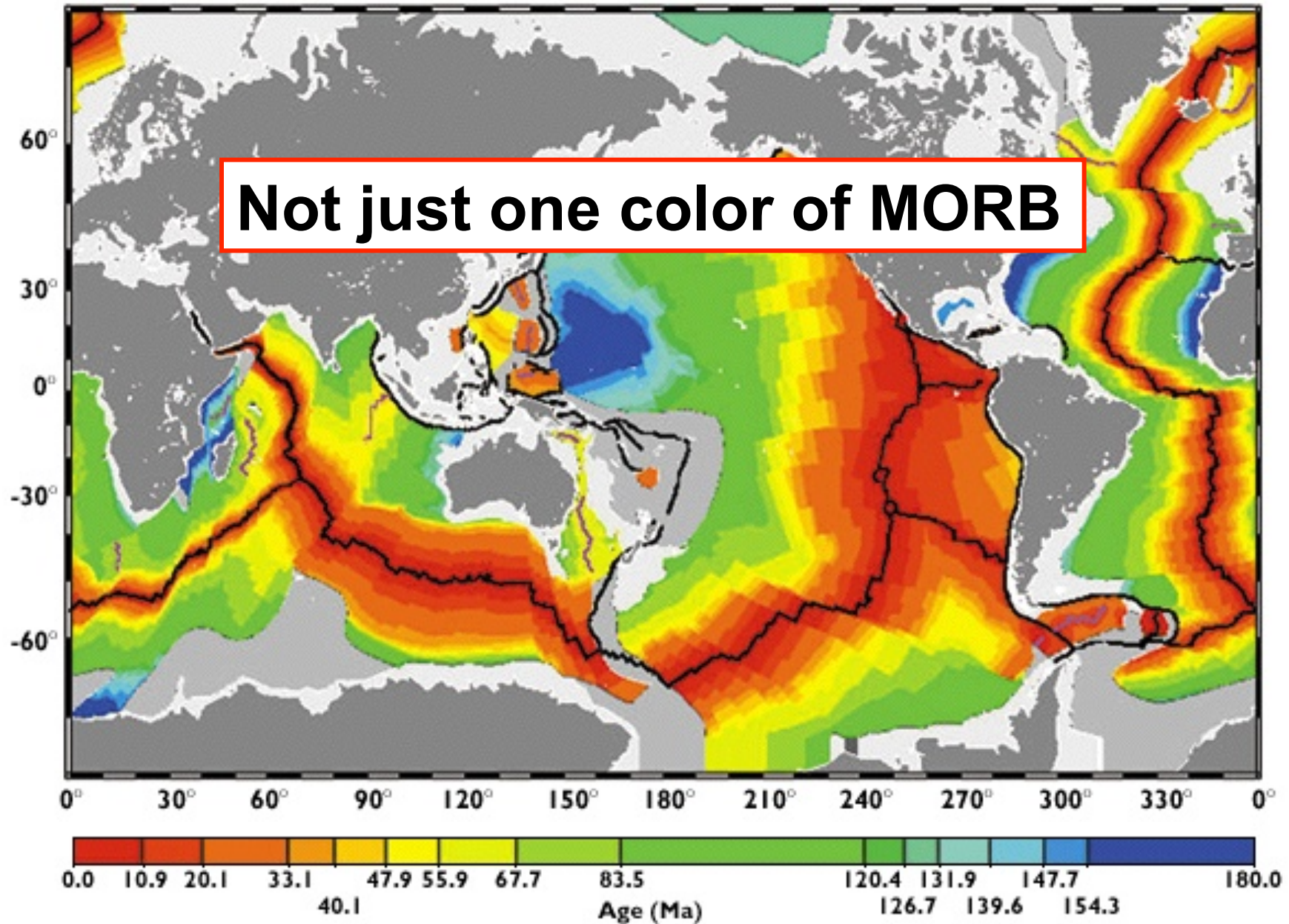
*Turcotte & Schubert (2002) 31 ng/g

Continental Crust
1300 ng/g U

Mantle
~12 ng/g U

*Chromatographic separation
Mantle melting & crust formation*

MORB: Distribution of ages at ridges



Spreading rate and structure

Fast East Pacific Rise



- Thermal structure - warmer
- Crust is thicker, lithosphere is thinner
- Higher degrees of melting
- Sustained magma chambers and volcanism
- **LESS COMPOSITIONAL DIVERSITY**

Slow Mid-Atlantic Ridge

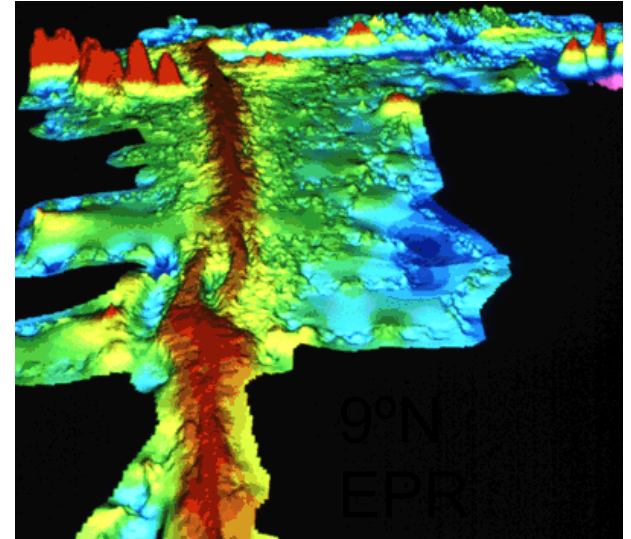


- Thermal structure - cooler
- Crust is thinner, lithosphere is thicker
- lower degrees of melting
- Episodic volcanism
- **GREATER COMPOSITIONAL DIVERSITY**

Highly incompatible elements in MORB

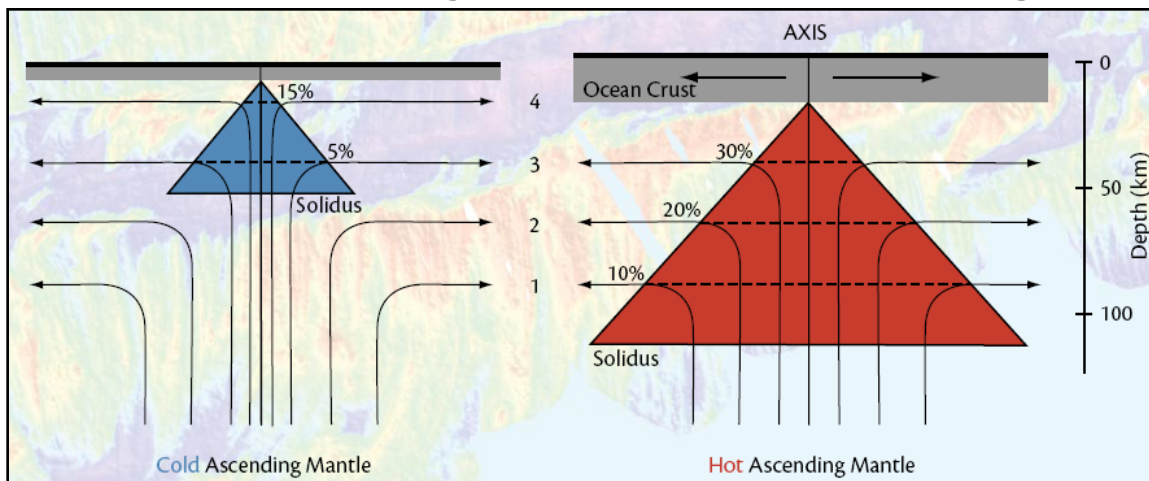
- MORB : active melting of the upper mantle
- production $20\text{-}25 \text{ km}^3 \cdot \text{yr}^{-1}$ (Crisp, 1984)
 - $\frac{3}{4}$ of all magma are intruded sills
- Negligible (crustal) contamination

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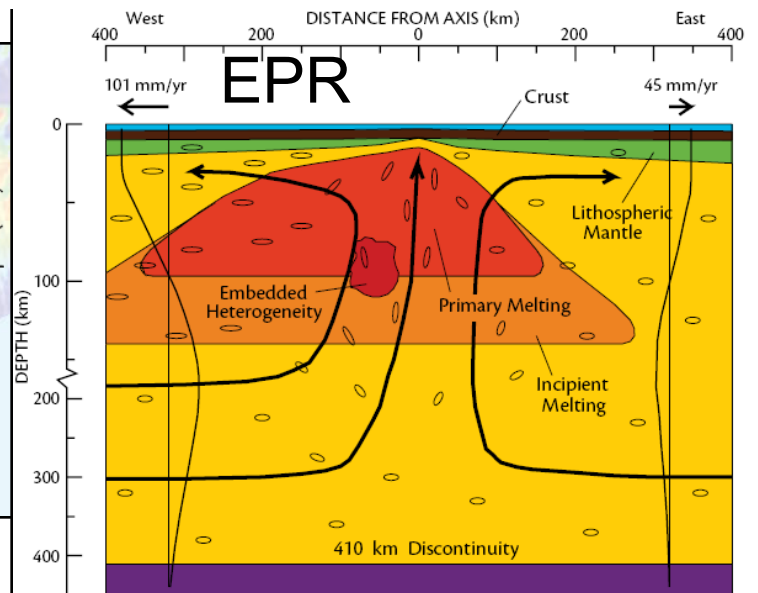
Cold upwelling

Hot upwelling



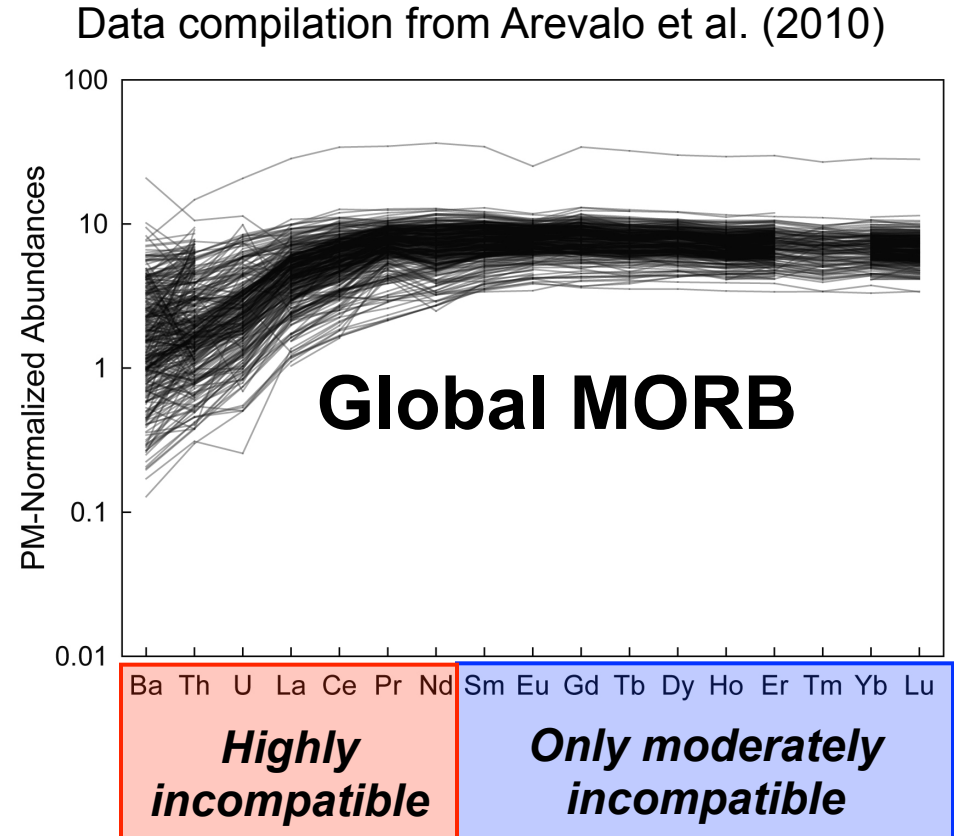
Langmuir & Forsyth (2007)

MELT seismic experiment

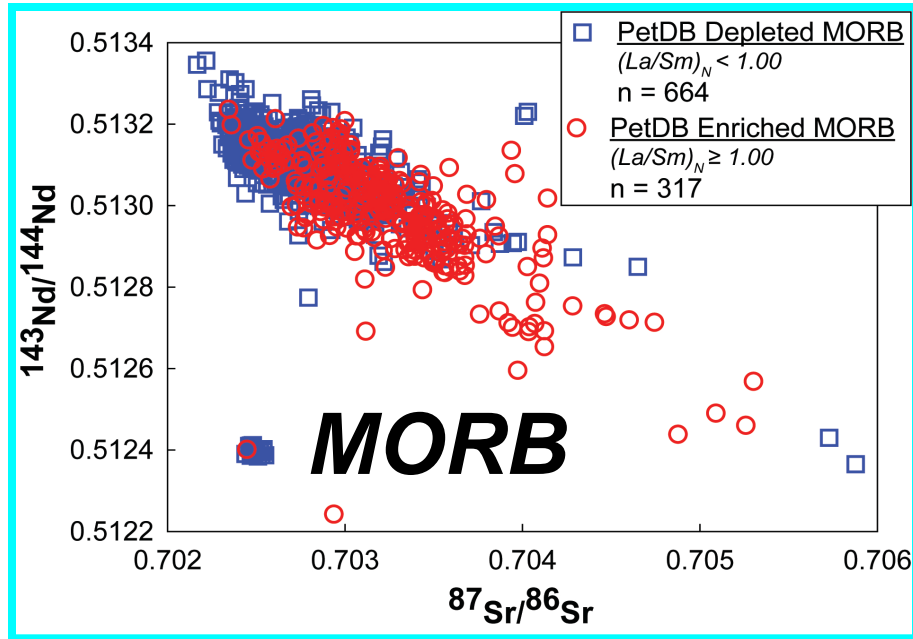


Highly incompatible elements

- Highly incompatible elements provide particularly sensitive tracers
 - Effectively removed from mantle sources at only minor degrees of partial melting
 - “See through” melting and crystallization
 - E.g., high-field strength elements (HFSE) and large-ion lithophile elements (LILE)

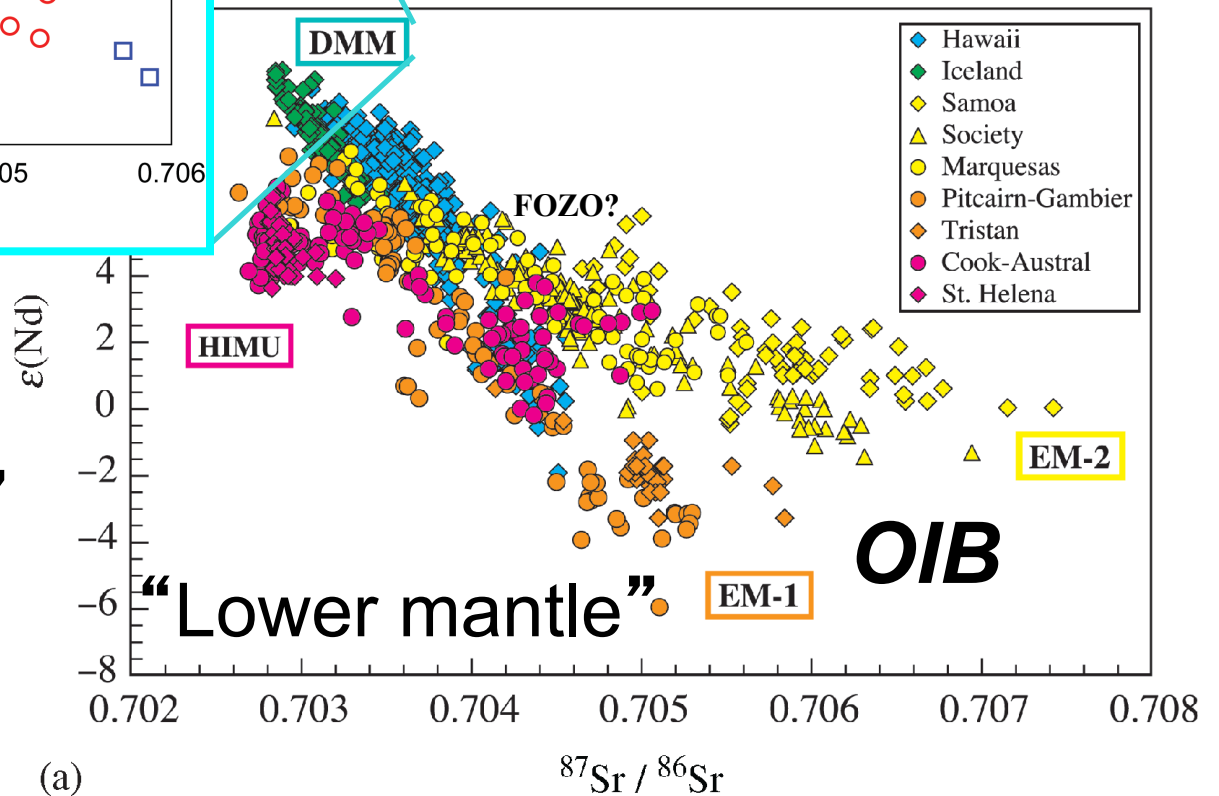


Mantle heterogeneity



- The modern mantle is heterogeneous in composition

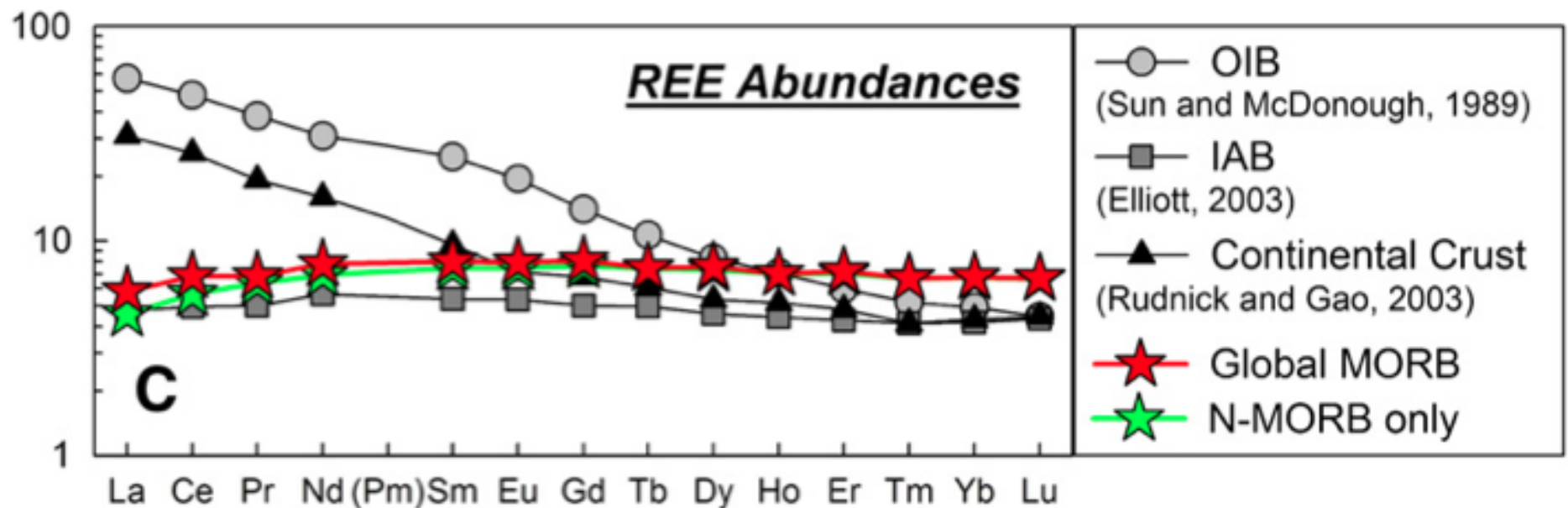
MORB is from the
 “Depleted Mantle”
 (DMM)



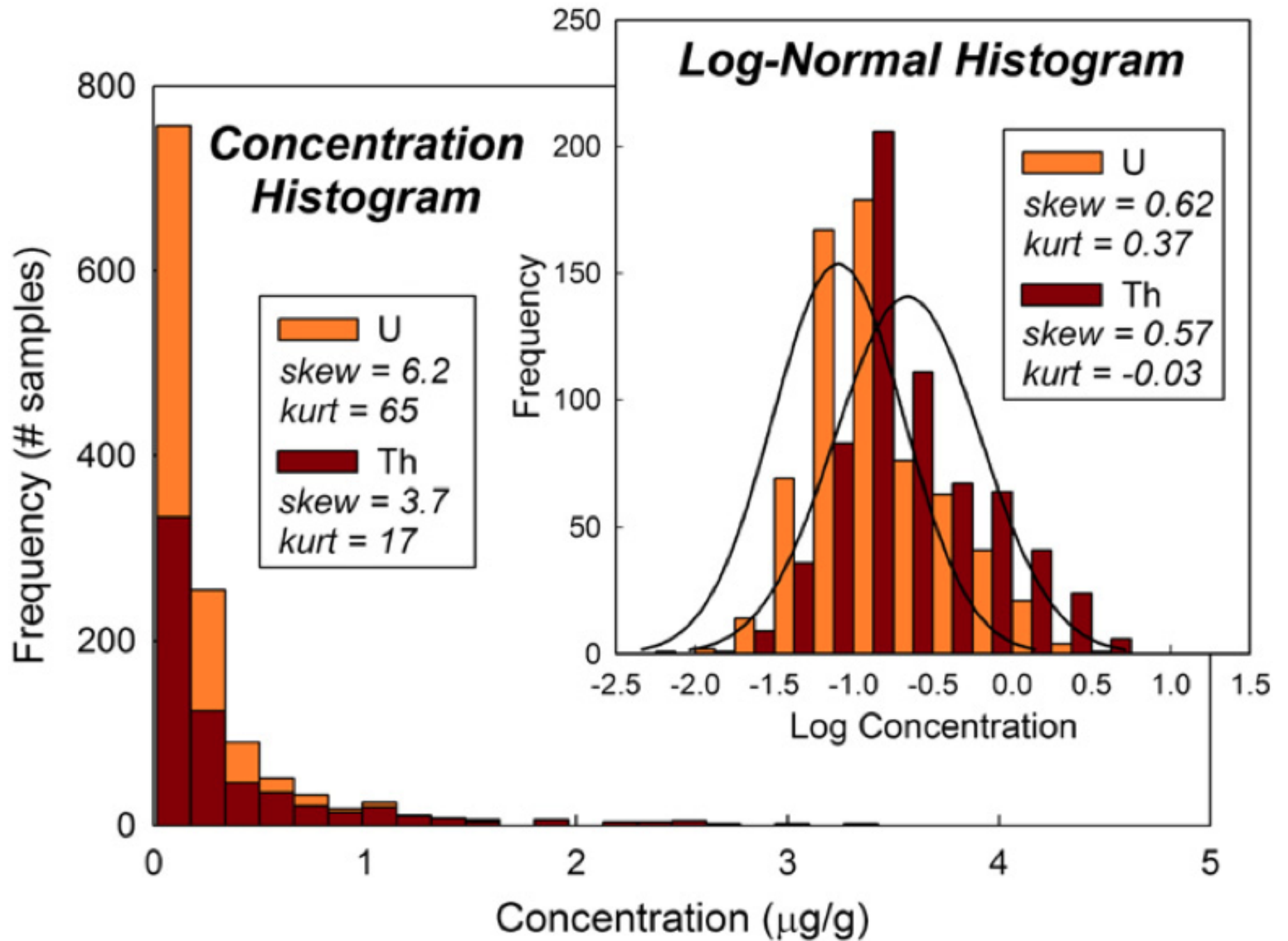
Hofmann (2003)

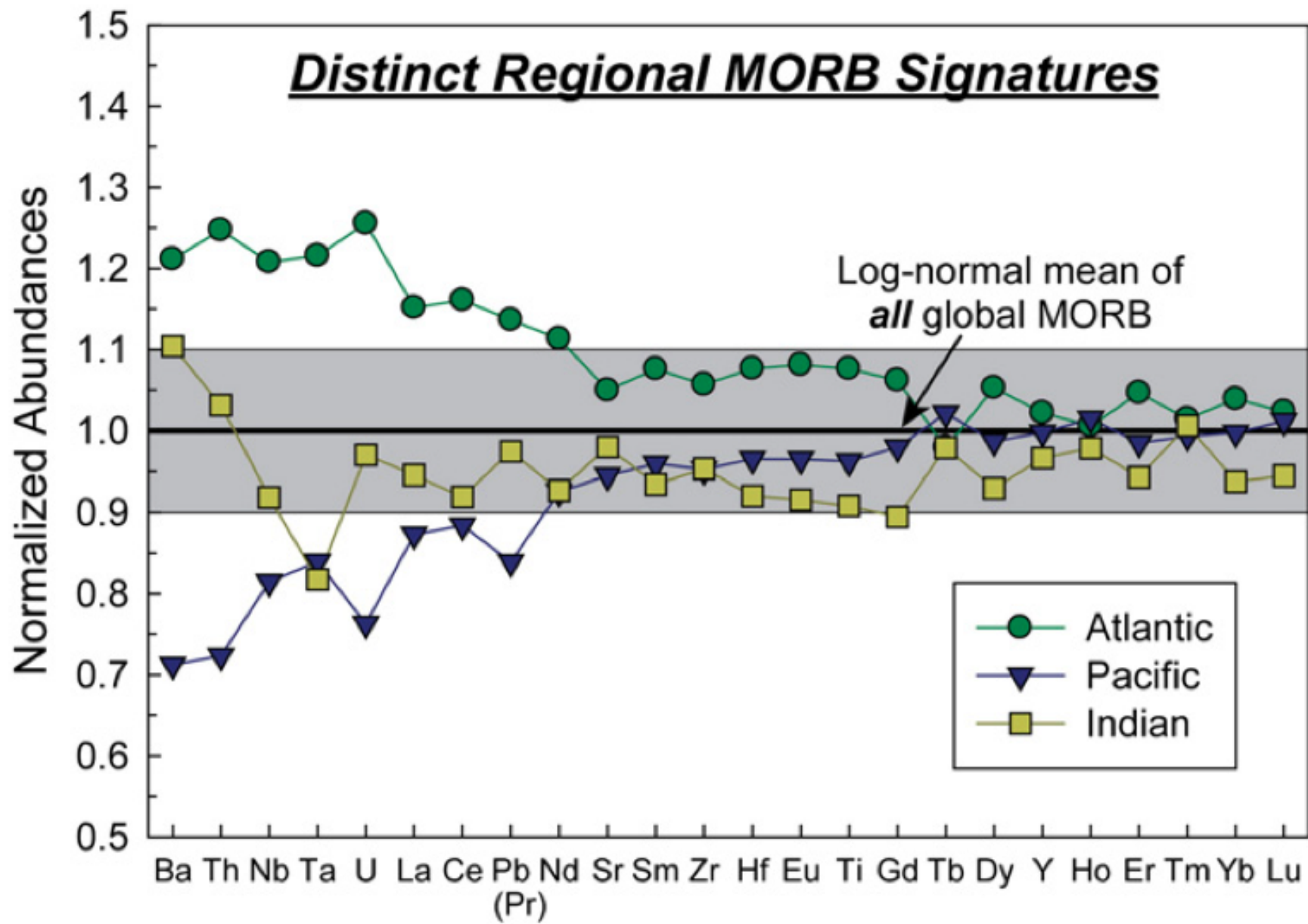
(a)

Complementary compositional nature of Continental Crust and **MORB** (and **OIB**)



Proportionally both should be added to make a **flat REE pattern for the BSE**





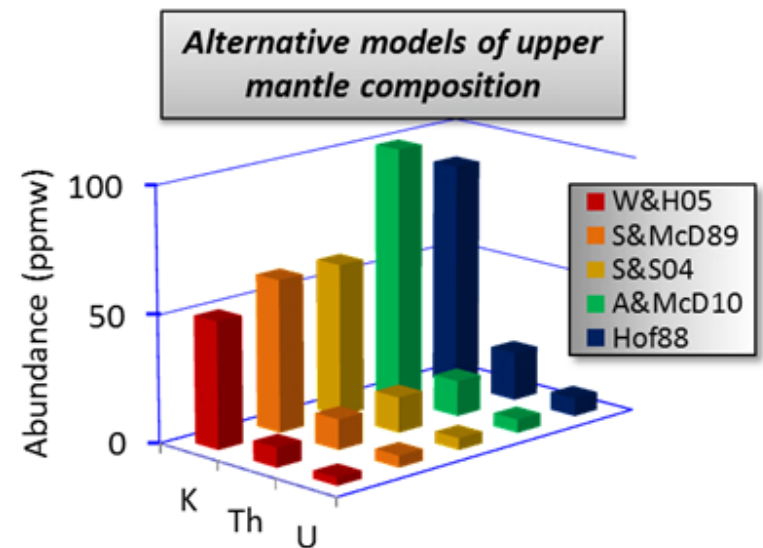
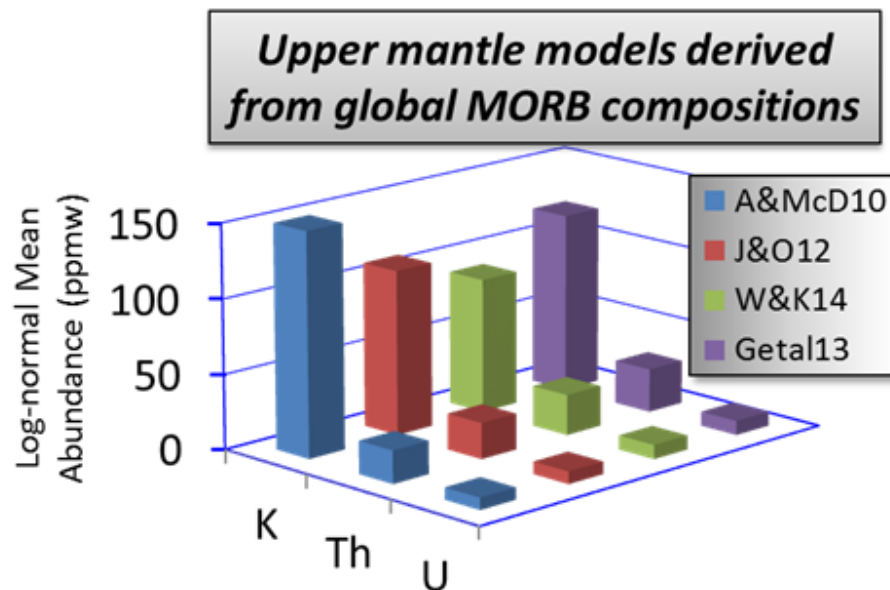
Modeling “Upper” Mantle Composition

Defining chemical composition for the upper mantle

N-MORB (normal-type): also referred as “depleted” type

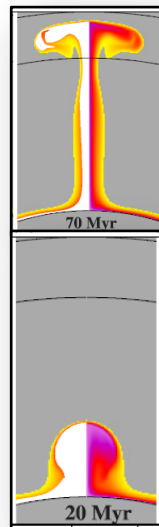
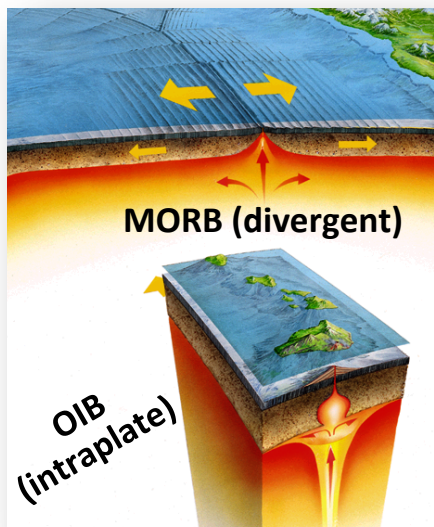
Often defined by $(\text{La}/\text{Sm})_N \leq 1.00$ or other chemical signature

Peridotite-derived models (melt residues)



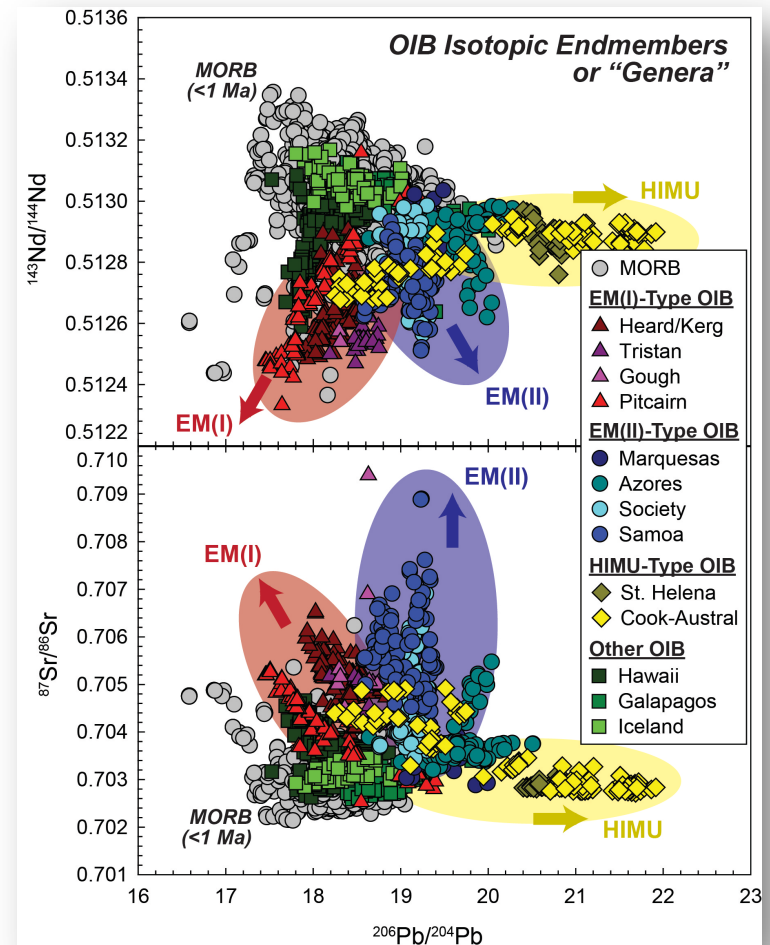
Modeling “Lower” Mantle Composition

- Ocean Island Basalts (OIBs) represent melting of deeper mantle sources
- Variable compositions reflect long-lived chemical heterogeneities, diverse melting conditions and rates of entrainment
- Enriched relative to MORB based on mass balance of elemental inventories and radiogenic heat production

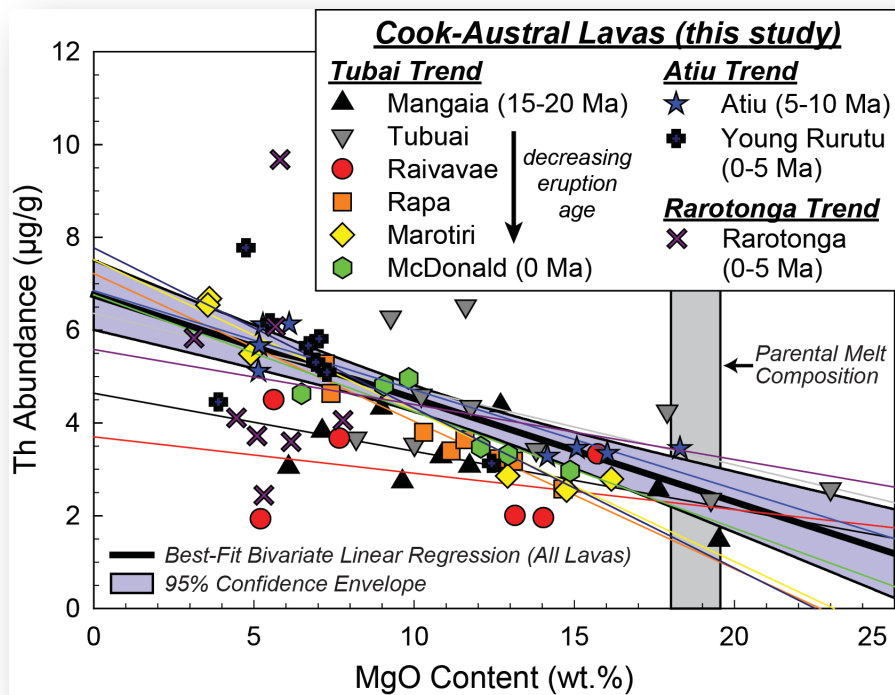


Lin & van Keken (2006) G^3

Data compiled by Stracke et al. (2013) G^3



A Case Study: The Cook-Australs

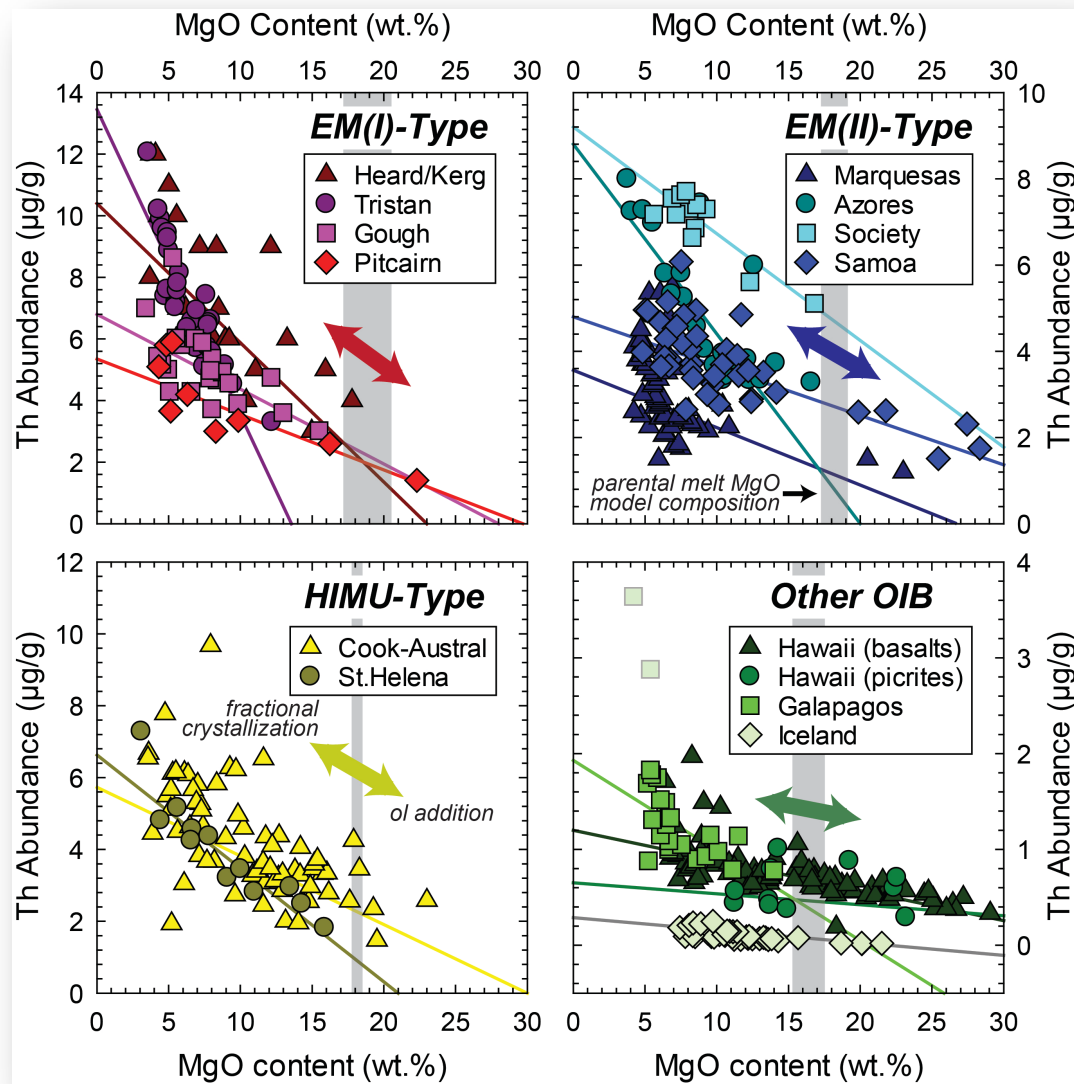


Arevalo et al. (2013) G^3

- Cook-Australs represent archetypal HIMU OIB
- Th is the most incompatible element (peridotite melting) that is neither fluid mobile (e.g., K or Rb) nor redox-sensitive (e.g., W or U)
- MgO maps the evolution of parental melt derivatives

We can calculate the abundance of Th in the near-primary parental melt(s) of the Cook-Australs through linear regression statistics (details forthcoming...).

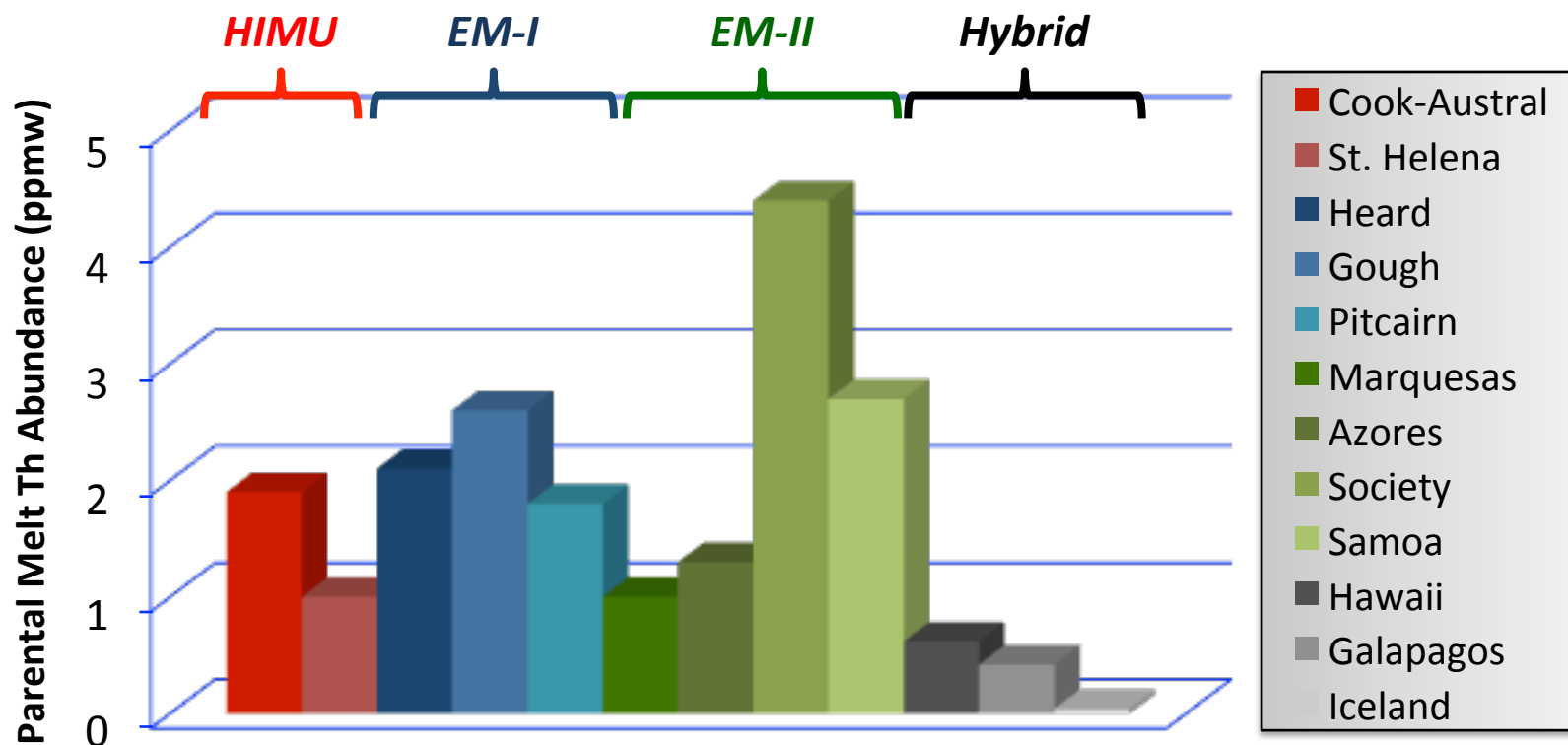
Universal Trends in Th versus MgO



Arevalo et al.
(2013) G^3

OIB Parental Melt Composition(s)

- Near-primary (or parental) melt compositions can be estimated by adding *ol* and *cpx* in constant proportions, dictated by $\text{CaO}/\text{Al}_2\text{O}_3$ relations, to equilibrate each individual lava with Fo_{90} olivine



OIB Source Composition(s)

- Assuming 5 – 10% partial melting of the source of the parental melts modeled here, consistent with studies of alkali OIB basalt petrogenesis
 - Corroborated via inverse modeling of Th/Nd and U/Sm
 - For comparison, MORB estimated to represent 8–15% melting

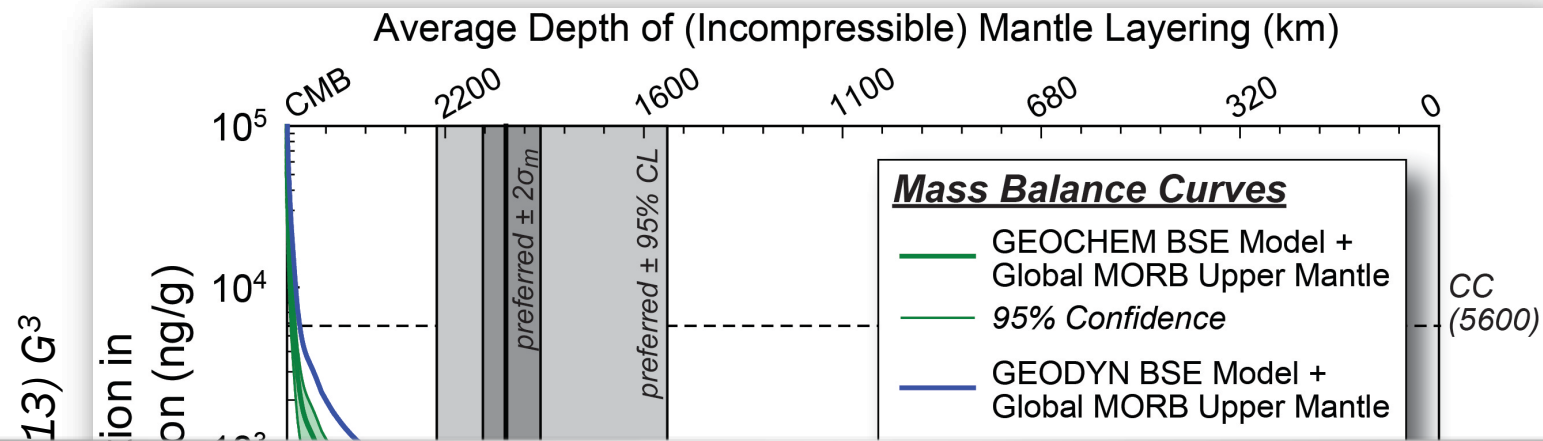
The OIB source region characterized by the samples considered here contains an average of:

(\geq) 160 ± 60 (± 20 , $2\sigma_m$) ng/g Th.

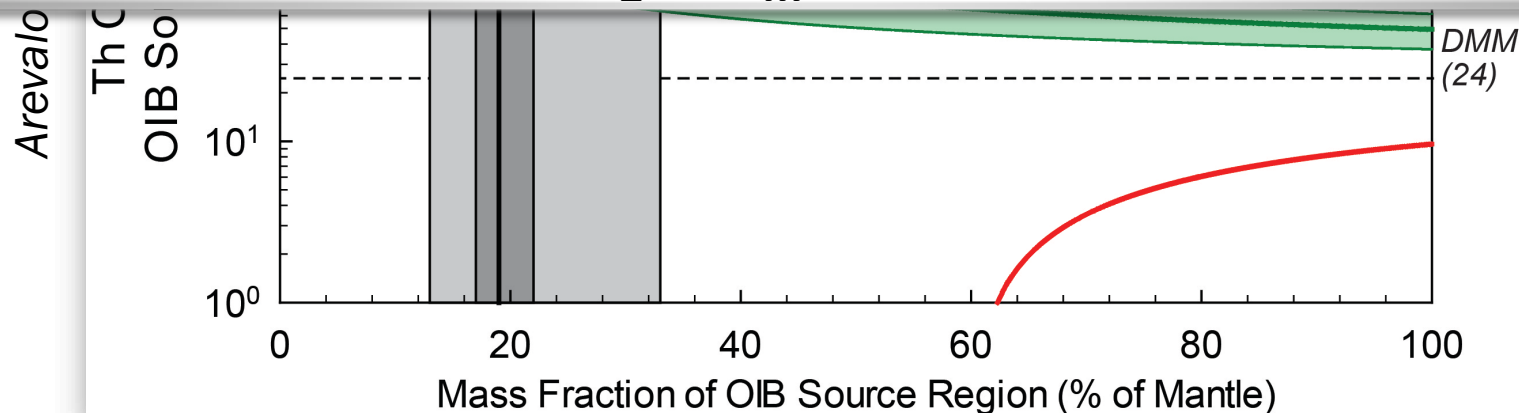
Assumptions include:

- 1) Lavas from each locality are cogenetic/representative of source;***
- 2) The bulk partition coefficient of Th is constant/well-constrained;***
- 3) Dilutional effects from entrainment not considered (for bounding);***
- 4) The source of melting is dominantly peridotitic (for bounding).***

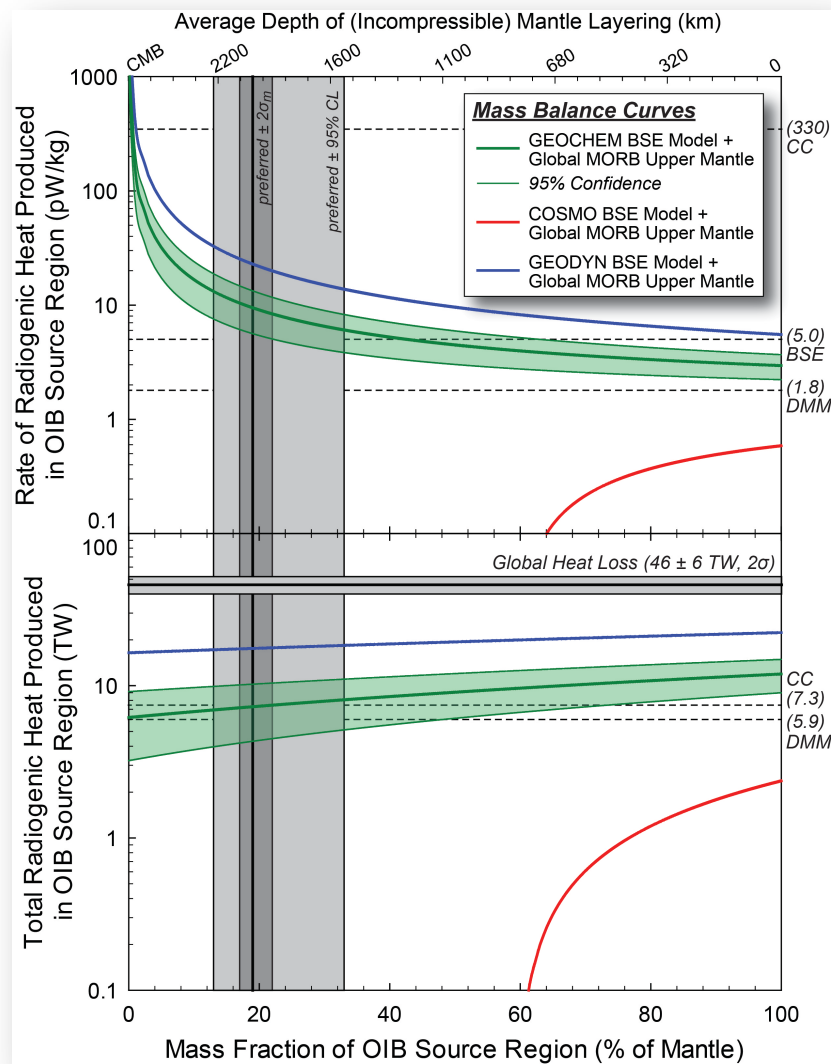
Size of the OIB Source(s)



Our “preferred” model suggests the OIB source region constitutes $(\leq)19^{+3}_{-2}$ ($2\sigma_m$)% of the mantle by mass.



Radiogenic Heat in the OIB Source(s)



- Our compositional model suggests the OIB source region produces:
 - (≥)9.5 pW/kg rate of heat generation
 - For comparison:
 - MORB source: <2 pW/kg
- The preferred mass balance curve (green) indicates the OIB source contributes:
 - (≤)7.3 TW radiogenic heat

